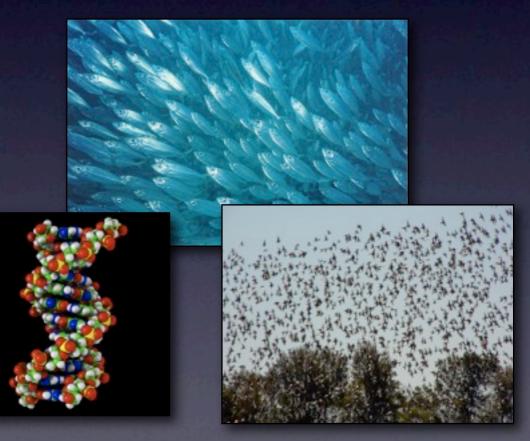
### Neutron Acceptance Diagram Shading Phil Bentley

# Why?

- New problems in optics often involve optimisation in manydimensional parameter spaces
- May have lots of local optima
- Analytical solutions might not exist

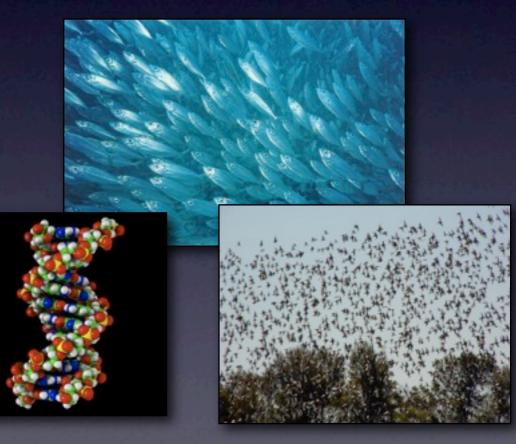
 Nature has solved such optimisation problems already



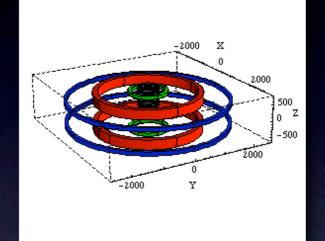
# Why?

- Particle swarms / genetic algorithms need 100 iterations or more to converge
- Each iteration evaluates 30-50 agents
- Therefore, 3000 simulations minimum

 Either use a cluster or something other than Monte-Carlo

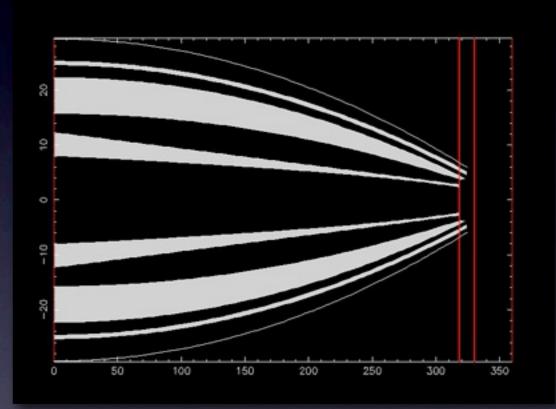


- Complex geometry with many coupled parameters
- SPAN-cryopad

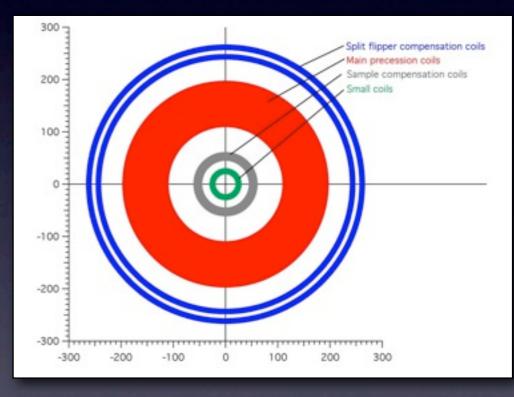


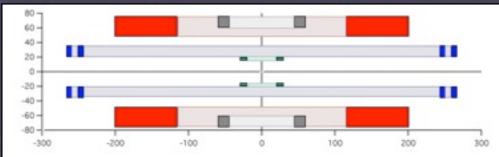
- multi-channel guides IN5
- WASP
- TYREX <sup>3</sup>He polariser (new model)

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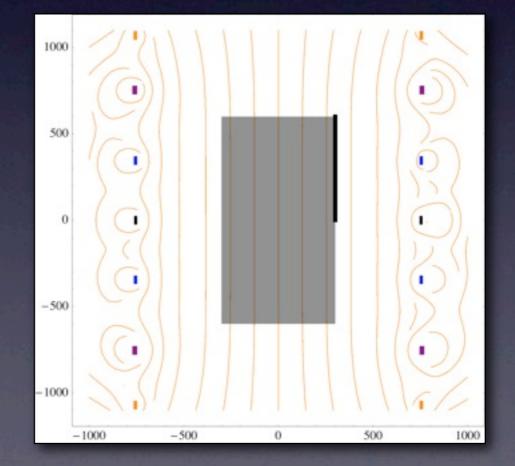


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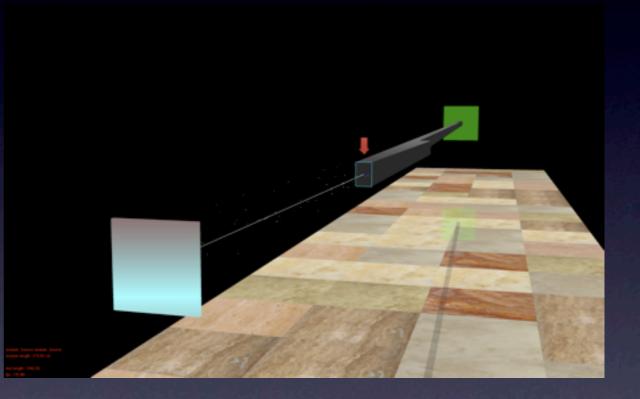


- Complex geometry with many coupled parameters
- SPAN-cryopad
- multi-channel guides IN5
- WASP
- TYREX <sup>3</sup>He polariser (new model)



### Full Optimisations

- Simulate the instrument all the way from moderator to detector
- All parameters free
- Don't care if parameters are strongly coupled
- Parameter limit: >47 but
   <650 with canonical algorithms</li>



- 100 000 000 trajectories entering the guide
- 666 seconds CPU time (11 minutes) => 74 trajectories hit the sample
- Statistical error = 12%
- 1% error needs 25 hours
- Optimisation of D22: 8-25 years (I cpu)

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# nads Speed Gain

1.5 x



http://flickr.com/photos/17393884@N00/5341048/ Flickr





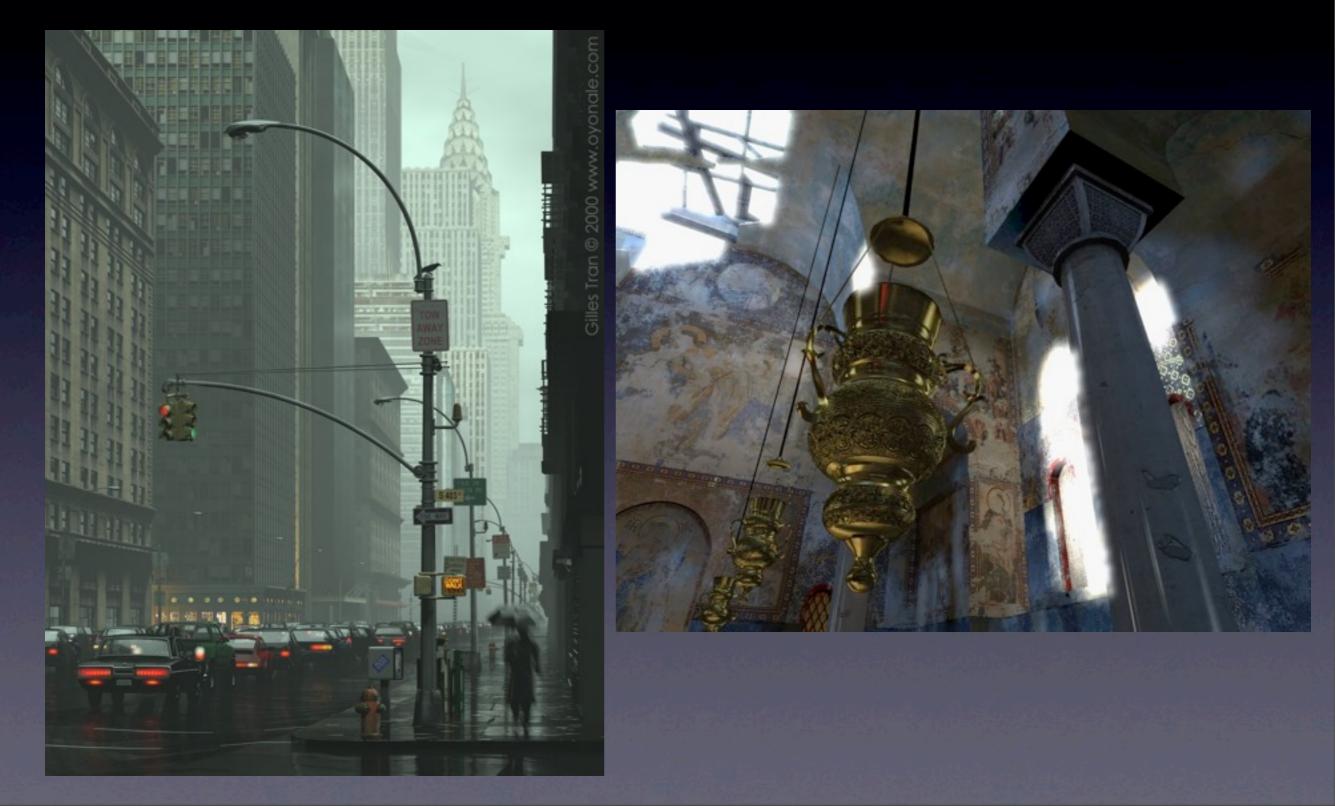
http://www.apolloarchive.com/apollo\_gallery.html

### Similar Problem Solved

- Ray tracing is routinely used in movies
- Very slow! Scenes rendered on huge clusters
- Vector graphics is close to ray trace quality
- 30 frames per second



### Similar Problem Solved

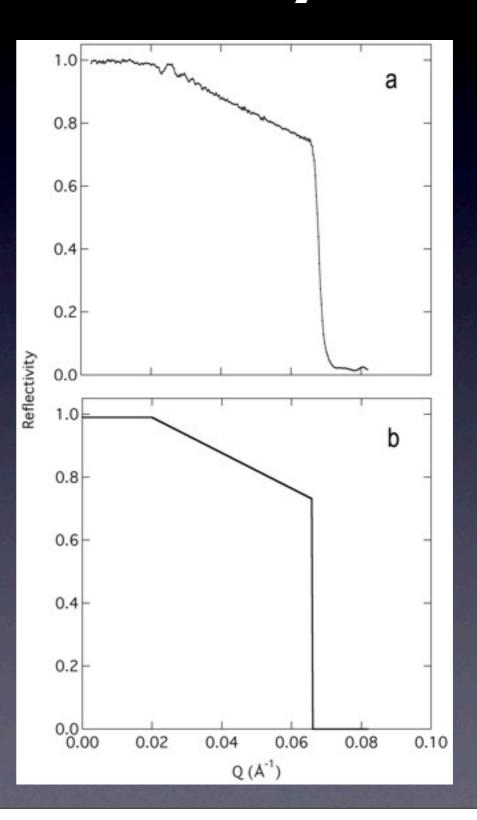


#### Neutron Bunches

- The analogy is to group trajectories into "similar" bunches and treat the group as one single object.
- Bunches are phase space regions with linear relation between trajectories
- How do we separate the bunches so the calculation remains accurate?

#### Guide Reflectivity

- Divide bunches along m=1 and critical m.
- Trajectories within a bunch are linearly dependent
- Guide reflectivity is an idealised curve
- R at m\_crit taken from Swiss neutronics



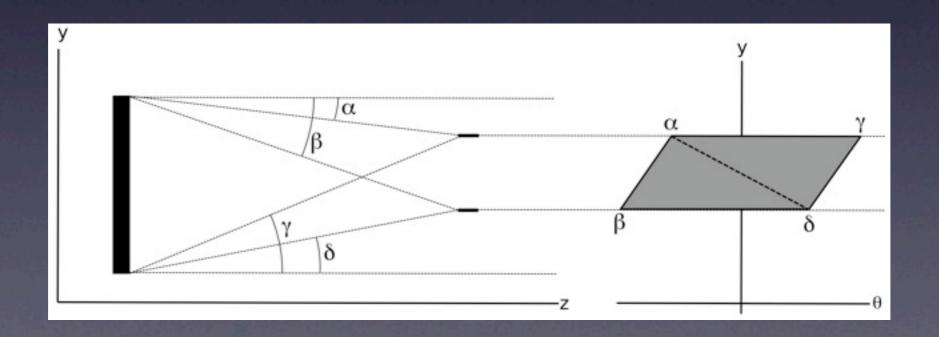
## Acceptance Diagrams

- Define boundaries in distance-divergence space
- Division is on module-by-module basis

#### Source Module

 Just like Monte-Carlo, we have a source plane and a virtual "exit window".  Exit window must be at least as large as subsequent module

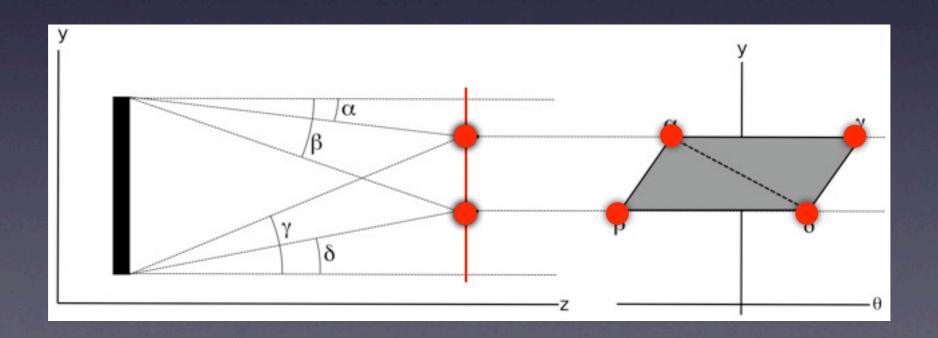
• Propagation = shear



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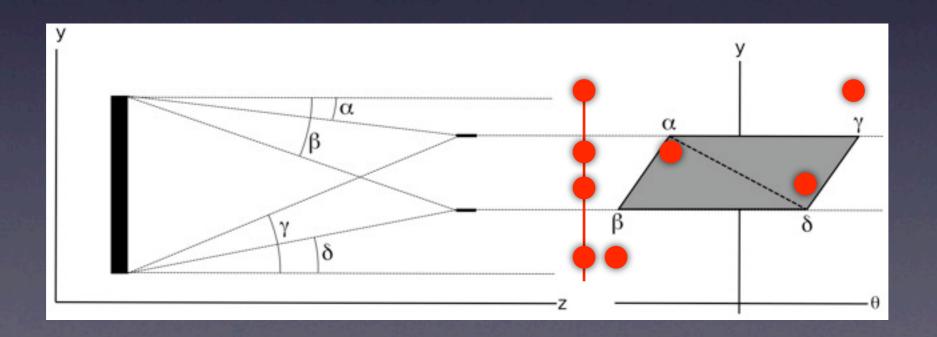
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#### Source Module

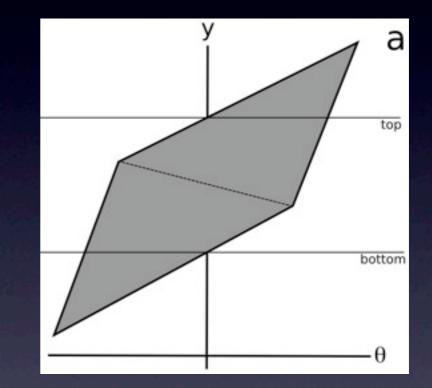
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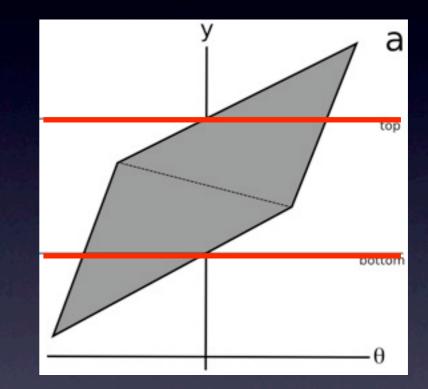
### Collimator/Aperture

- For aperture clip off extrema in space axis
- For collimator, clip off extrema in divergence axis



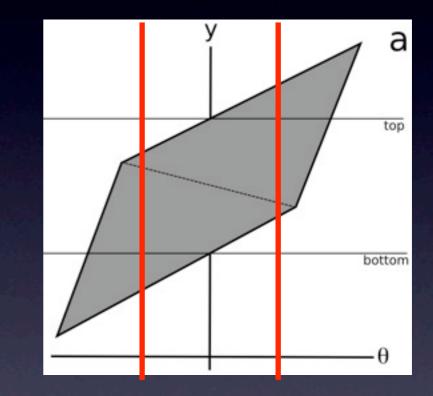
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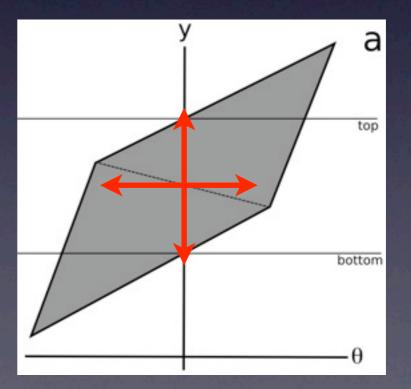
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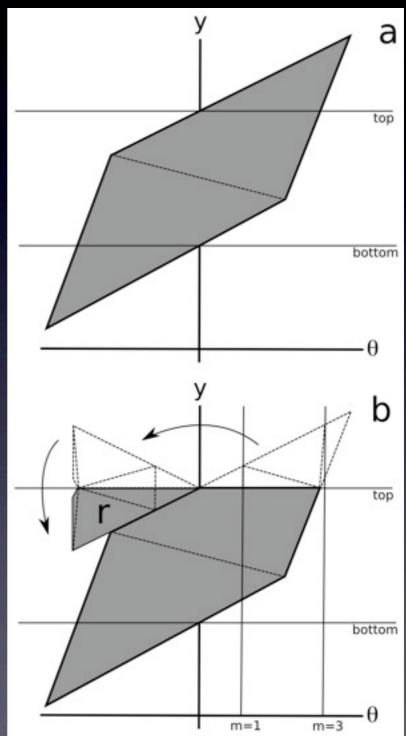
## Change Coordinates

- Rotate beam axis
- Translate beam axis
- Both are a translation of phase space volumes



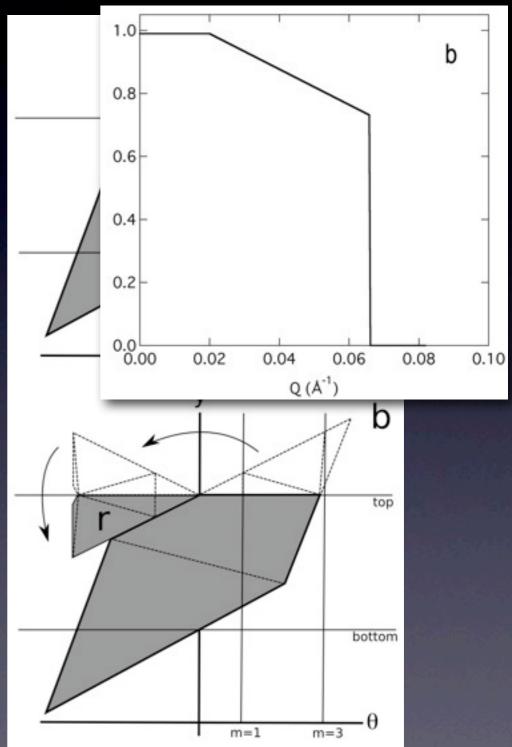
### Guide Module

- Kill neutrons that miss the entrance
- Propagate neutrons to the end of the guide
- Divide along m boundaries
- Multiply reflected weight by reflectivity



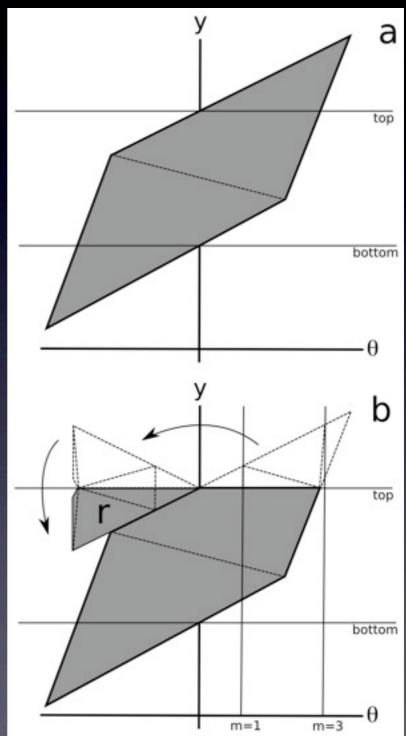
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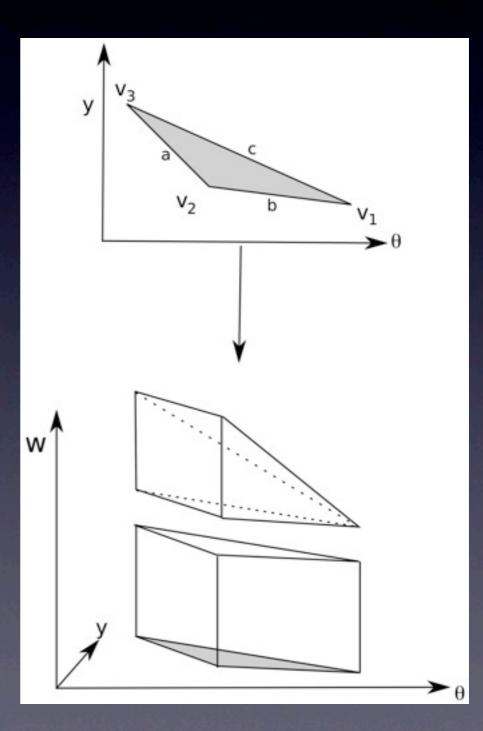
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### Statistical Weight

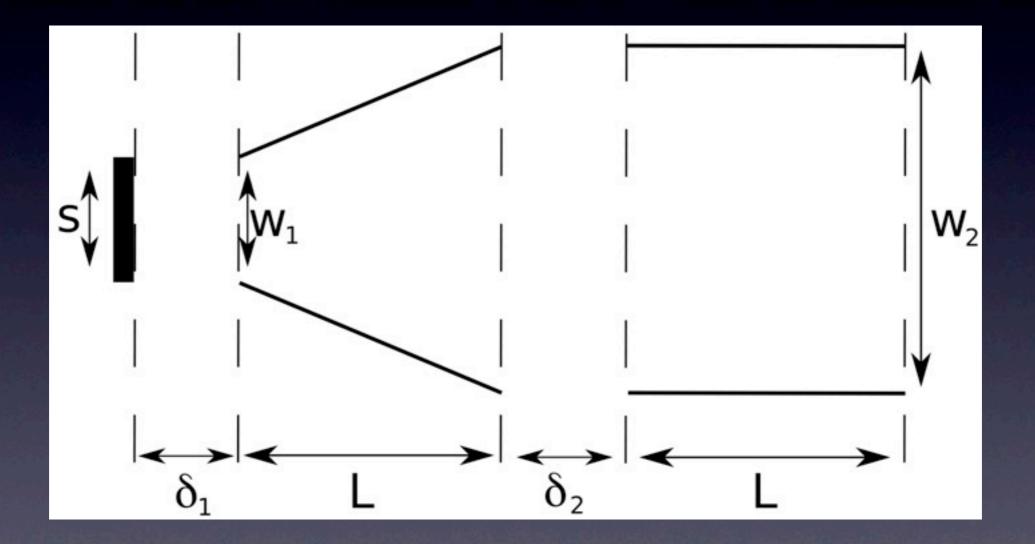
- Use triangular primitives
- Statistical weight of a triangle is a volume calculation
- Right-wedge plus square-based pyramid
- Only difficulty is very thin triangles (known problem in CS)

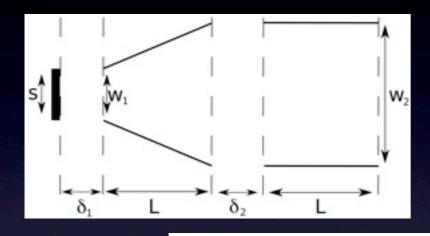


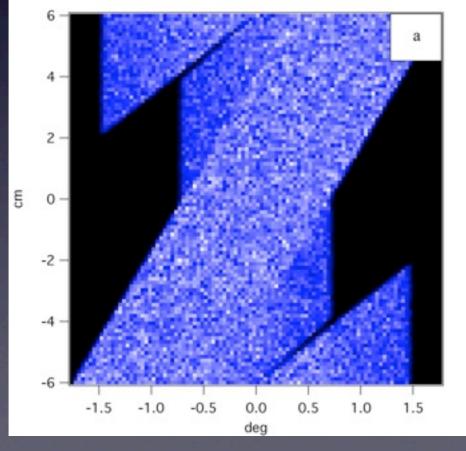
## Guides Supported

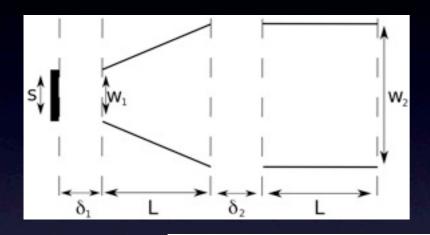
- Straight
- Curved parallel
- Converging
- Diverging
- Arbitrary

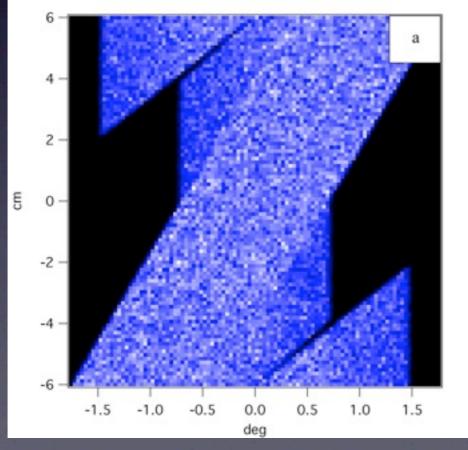


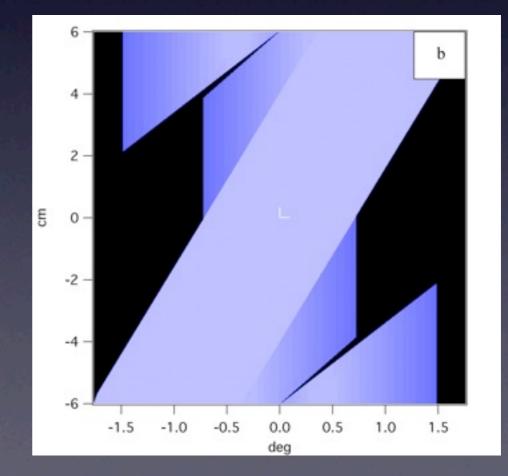


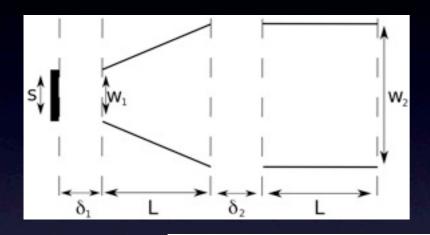


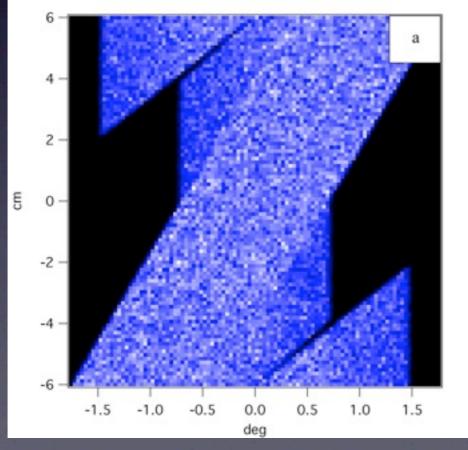


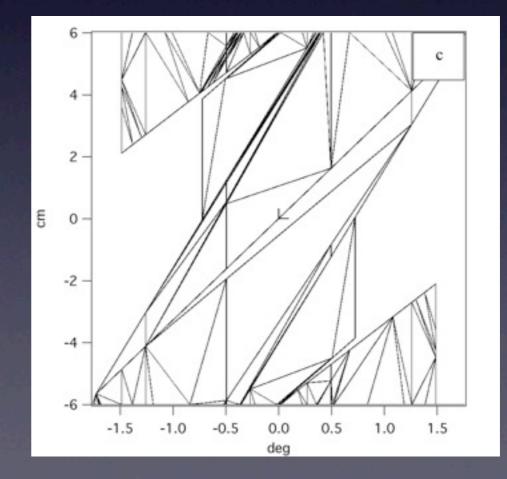


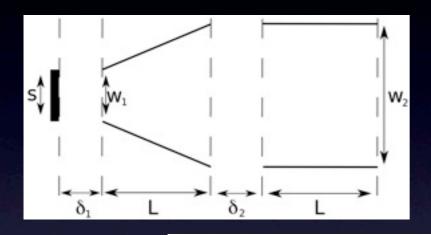


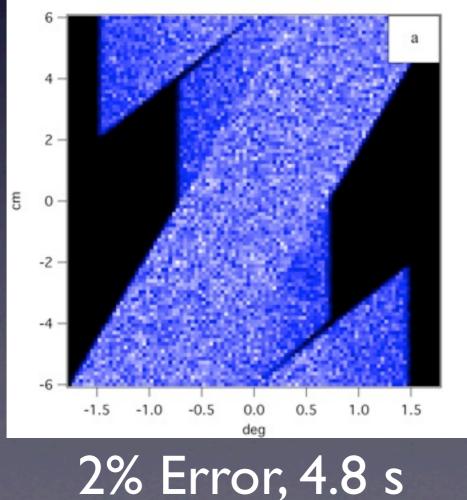


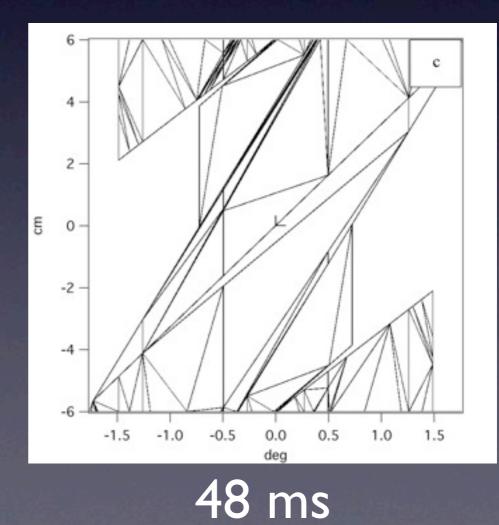






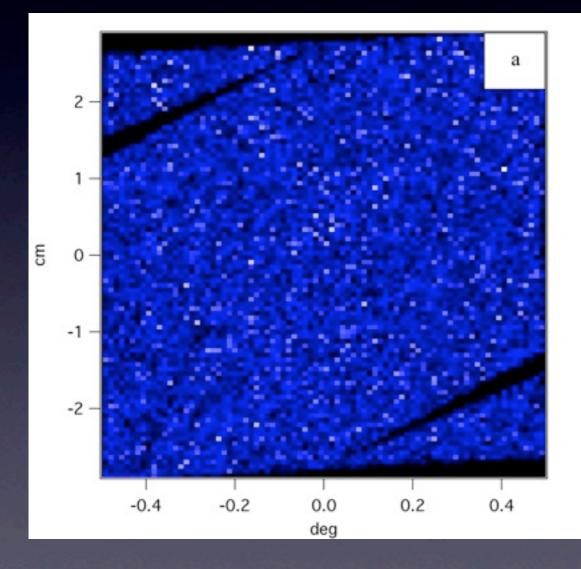




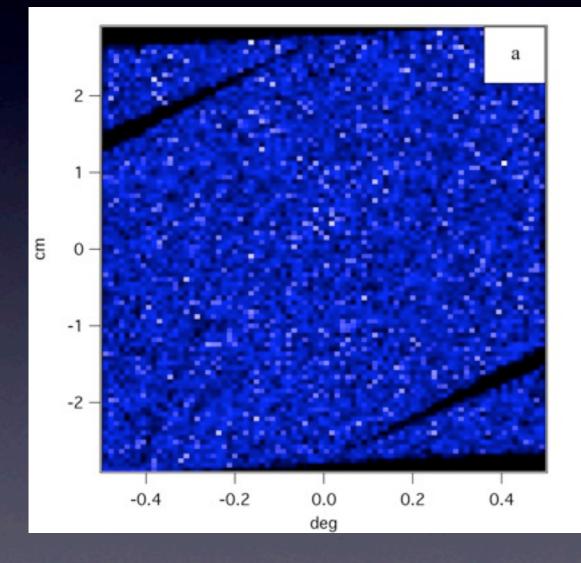


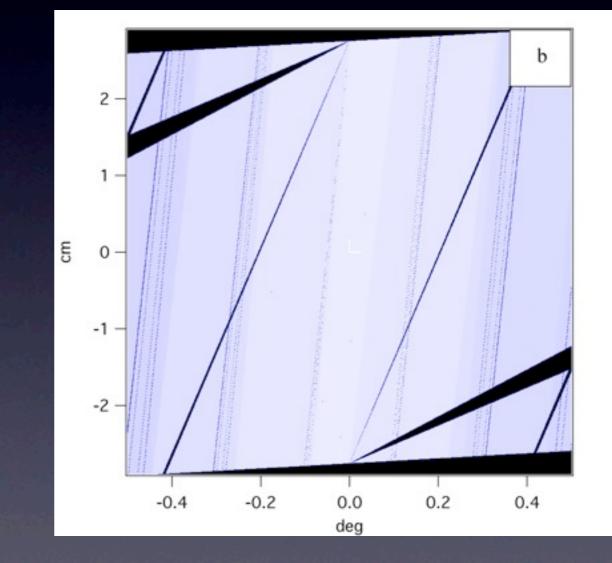
Monday, 17 August 2009

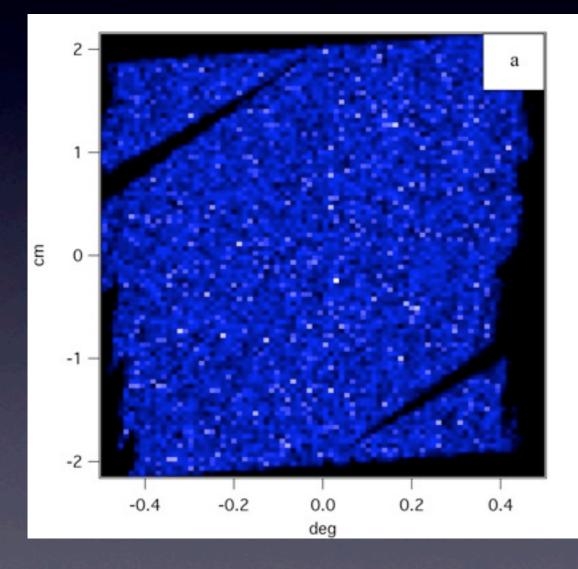
# Test 2: D22 (V)



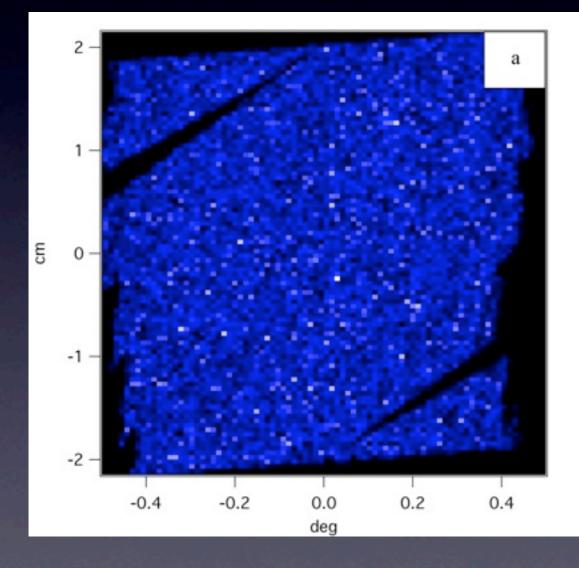
Monday, 17 August 2009

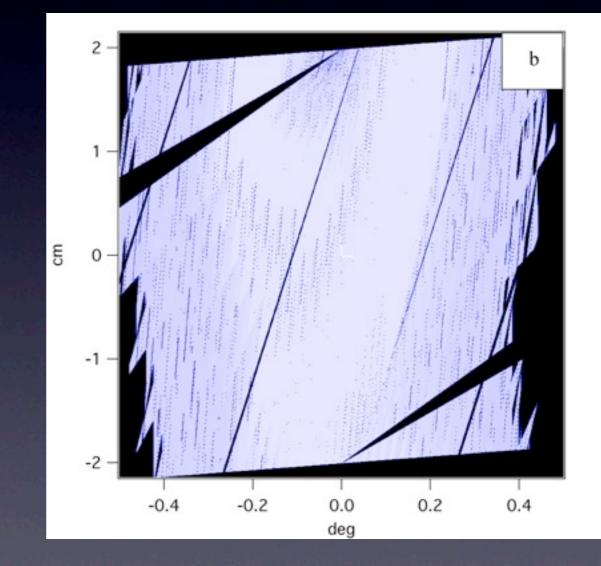


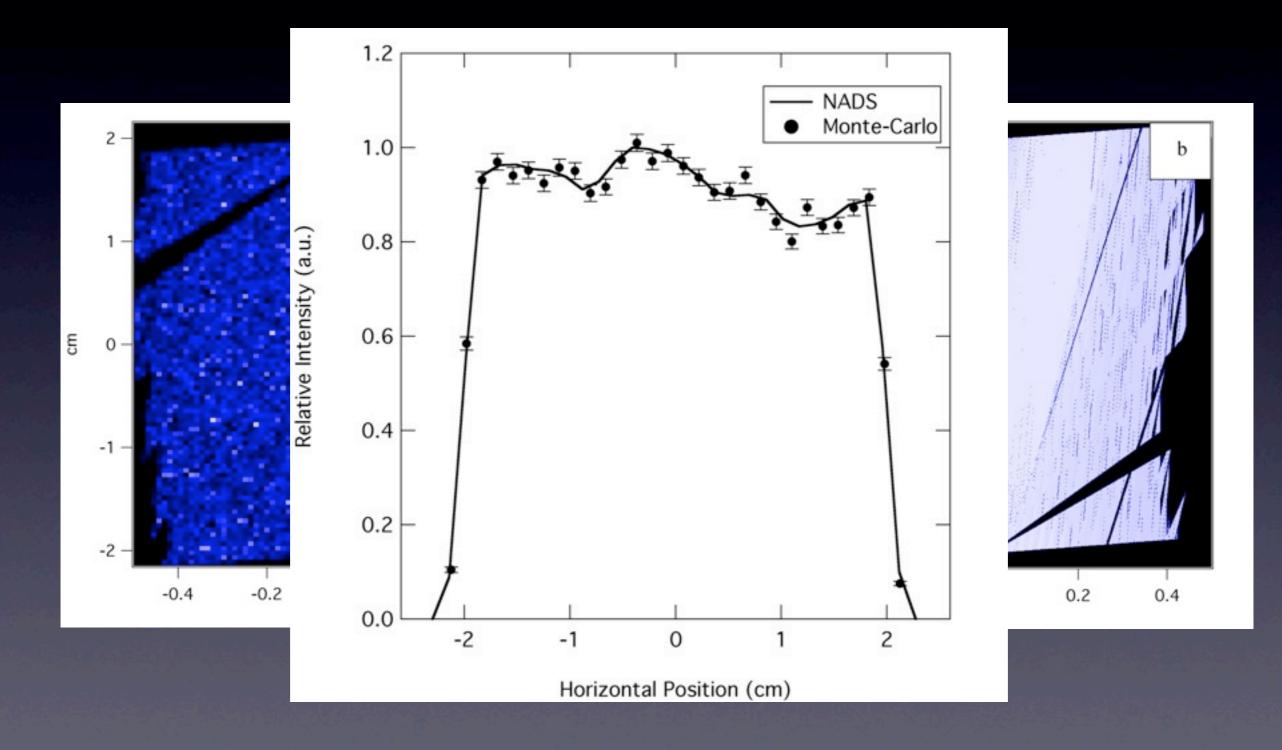


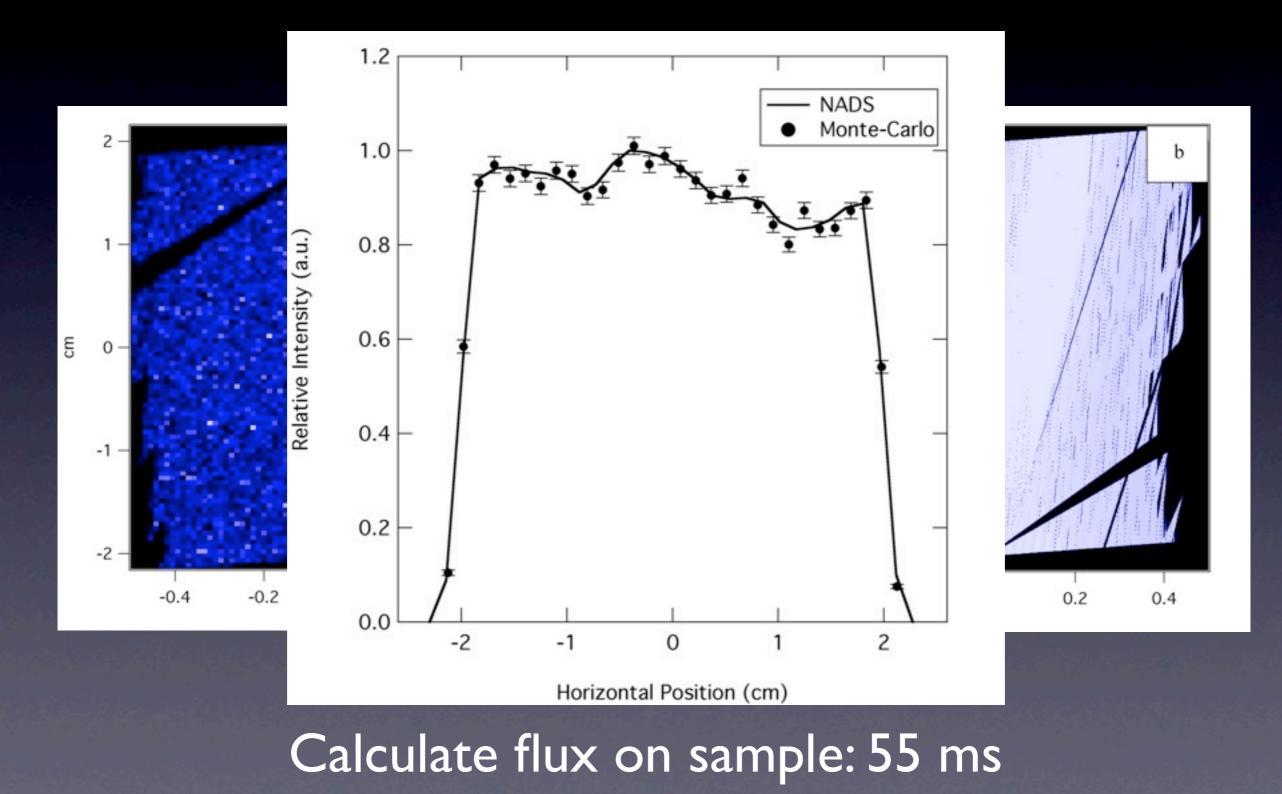


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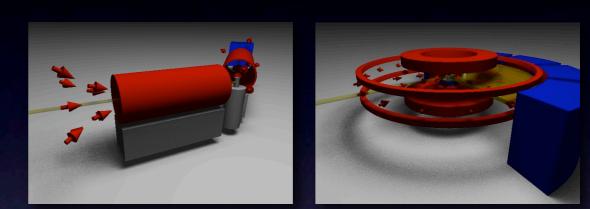




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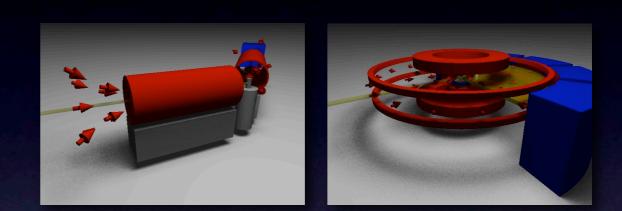
## Test 3:WASP

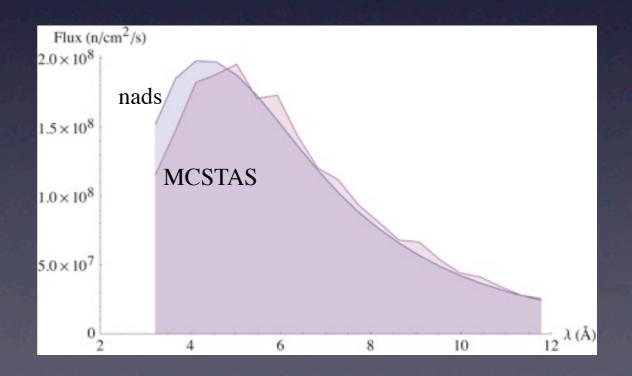
- Agreement between NADS and MCSTAS on sample
- MCSTAS model Independently calculated by Peter Fouquet from moderator to Sample
- ~5 seconds for nads per wavelength (3 mins for white beam)



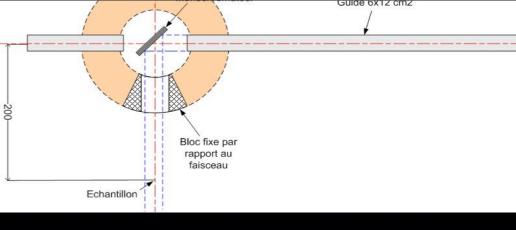
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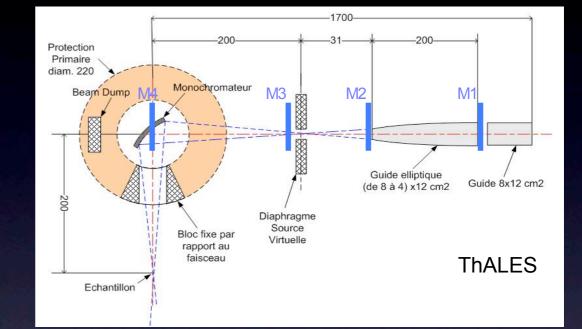
# Test 4:



• Agreement between SIMRES / RESTRAX and nads  $I = \int_{0.7}^{10} (\frac{dI}{d\lambda}) d\lambda$ 

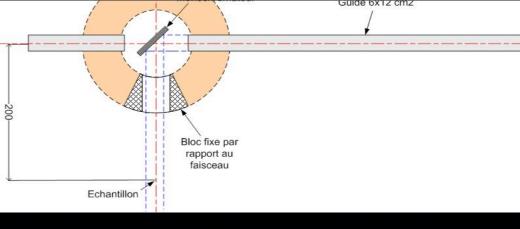
• Trace  $= \int_{0.7}^{10} (\frac{d\phi}{d\lambda}) d\lambda$  virtual source (M3)

 $\frac{dI}{d\lambda}\Delta\lambda$ 



 ~40 ms for nads per wavelength (1.7 sec for white beam)

# Test 4:

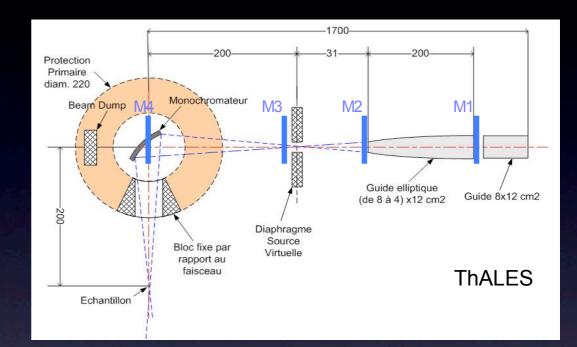


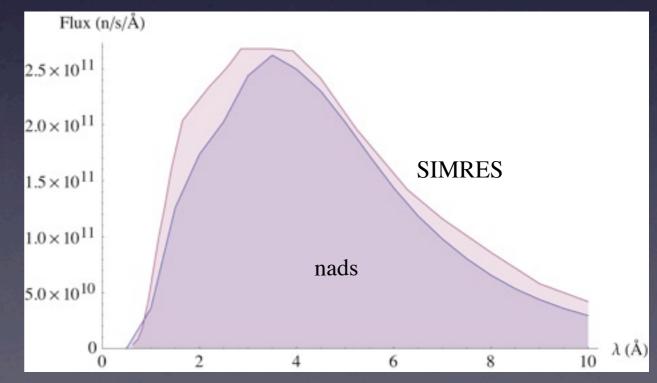
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 $\frac{dI}{d\lambda}\Delta\lambda$ 

 ~40 ms for nads per wavelength (1.7 sec for white beam)





#### For Coders

- nads kernel in C++
- Understands xml input from file or pipe
- xml schema lets you check syntax of handcoded simulations
- Uses getopt with varying degrees of verbosity

- nads -q prints only the output flux
- Fully scriptable using bash / c / mathematica...

```
<module>
    <rotation>
        <elemname>New Rotation Module</elemname>
        <hRotationAngle>0.6</hRotationAngle>
        <vRotationAngle>0</vRotationAngle>
    </rotation>
</module>
<module>
    <convergingGuide>
        <elemname>Separator</elemname>
        <width>8</width>
        <height>12</height>
        <exitWidth>18</exitWidth>
        <exitHeight>12</exitHeight>
        <length>2.77</length>
        <mnumber>3</mnumber>
    </convergingGuide>
</module>
```

## Live Demonstration

#### • GUI

- 3d instrument visualisation with OpenGL
- Mathematica scripts

#### Limitations

- Each nads calculation is monochromatic (white beam possible via looping)
- Small angle approximation
- Separable vertical & horizontal channels
- Round objects approximated by rectangular objects of equal area (round sources etc)

## Acknowledgements

#### • ESFRI

- Ken Andersen
- Klaus Habicht (HMI / BLZ)
- Leo Cussen (Cussen Consulting)